

2.2. MEAN FREE PATH IS GIVEN BY  $\ell = \frac{1}{\sigma n}$   
 CROSS SECTION,  $\sigma = \pi r^2$

$$\ell = \frac{1}{\pi r^2 n}$$

$$\ell_{\star} = \frac{1}{\pi r_{\star}^2 n_{\star}} = [6.2 \times 10^{17} \text{ Mpc}]$$

$$\ell_g = \frac{1}{\pi r_g^2 n_g} = [8.0 \times 10^4 \text{ Mpc}]$$

2.4 WE KNOW MASSES ARE NON-NEGATIVE  
 ⇒ SET LIGHTEST NEUTRINO TO  $[m_{\nu_e} = 0]$

$$\text{then } m_{\nu_\mu} = \sqrt{5 \times 10^{-5} \text{ eV}^2/c^4} = [.0071 \text{ eV}/c^2]$$

$$m_{\nu_\tau} = \sqrt{3 \times 10^{-3} \text{ eV}^2/c^4 + m_{\nu_\mu}^2} = [.055 \text{ eV}/c^2]$$

2.5  $\frac{dE}{F} = -Kdr$

$$\ln E = -Kr + C$$

$$E = Ce^{-Kr} \quad \leftarrow \quad F = h\nu = h\left(\frac{c}{\lambda}\right)$$

$$\frac{1}{\lambda} = Ce^{-Kr} \quad (\text{RESCALING OVERALL CONSTANT})$$

$$z = \frac{\lambda_{OB} - \lambda_{EM}}{\lambda_{EM}}$$

$$\leftarrow \quad \lambda_{EM} = 0$$

$$= Ce^{-K\lambda_{EM}} \left( C^{-1}e^{K\lambda_{OB}} - C^{-1}e^{K\lambda_{EM}} \right)$$

$$= e^{K\lambda_{OB}} - 1$$

for  $z \ll 1$  we can Taylor expand RHS

$$z = Kr + O((Kr_{OB})^2)$$

where  $Kr_{OB} \ll 1$

$$\text{if } z = \frac{h\nu}{c} r, \quad K = \frac{h\nu}{c}$$

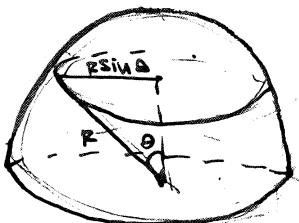
3.2 (3.9)  $ds^2 = dr^2 + R^2 \sin^2 r/R d\theta^2$

$$\Rightarrow \boxed{d\theta = \frac{ds}{R \sin r/R}}$$

As  $r \rightarrow \pi R$ ,  $d\theta \rightarrow \infty$

THIS CORRESPONDS TO ANTIPODAL POINTS  
ON THE SPHERE WHERE  $\theta$  IS NOT WELL DEFINED

3.3



$$\Delta = r/R$$

$$\Rightarrow C = 2\pi R \sin r/R$$

$$|C_e - C| > 1 \text{ m}$$

$$|2\pi r - 2\pi R \sin r/R| > 1 \text{ m}$$

$$R \sin r/R < r - \frac{1}{2\pi}$$

$$\boxed{r > 38 \text{ km}}$$

3.4 a) for  $K=+1$ , the surface is compact  $\Rightarrow$   $\exists$  maximum area

$$A = (\text{AREA OF SURFACE}) = \boxed{4\pi R^2}$$

(IMAGINE AN ARBITRARILY SMALL TRIANGLE ON THE SURFACE,  
NOTE THAT THE COMPLEMENT OF THIS TRIANGLE IS ALSO  
A TRIANGLE!)

b) for  $K=0$   $\boxed{A = \infty}$

c) (3.10)  $3\gamma = \pi - A/R^2$   $\gamma = \text{ANGLE OF TRIANGLE (EQUILATERAL)}$

$$\gamma \geq 0 \Rightarrow A = \boxed{\pi R^2}$$

$$\begin{aligned}
 3.5 \quad dx &= \sin\theta \cos\phi dr + r \cos\theta \cos\phi d\theta + r \sin\theta (-\sin\phi) d\phi \\
 dx^2 &= \sin^2\theta \cos^2\phi dr^2 + r \sin\theta \cos\theta \cos^2\phi dr d\theta - r \sin^2\theta \cos\theta \sin\phi dr d\phi \\
 &\quad + r \sin\theta \cos\theta \cos^2\phi d\theta^2 + r^2 \cos^2\theta \cos^2\phi d\theta^2 - r^2 \cos\theta \sin\theta \cos\theta \sin\phi d\theta d\phi \\
 &\quad - r \sin^2\theta \cos\theta \cos^2\phi d\phi^2 - r^2 \cos\theta \sin\theta \cos\theta \sin\phi d\phi^2 + r^2 \sin^2\theta \sin^2\phi d\phi^2 \\
 &= \sin^2\theta \cos^2\phi dr^2 + 2r \sin\theta \cos\theta \cos^2\phi dr d\theta - 2r \sin^2\theta \cos\theta \sin\phi dr d\phi \\
 &\quad + r^2 \cos^2\theta \cos^2\phi d\theta^2 - 2r^2 \cos\theta \sin\theta \cos\theta \sin\phi d\theta d\phi \\
 &\quad + r^2 \sin^2\theta \sin^2\phi d\phi^2 \\
 dy &= \sin\theta \sin\phi dr + r \cos\theta \sin\phi d\theta + r \sin\theta \cos\phi d\phi \\
 dy^2 &= \sin^2\theta \sin^2\phi dr^2 + r \sin\theta \cos\theta \sin\phi dr d\phi + r \sin^2\theta \cos\theta \sin\phi dr d\phi \\
 &\quad (\text{symmetric}) + r^2 \cos^2\theta \sin^2\phi d\theta^2 + r^2 \cos\theta \sin\theta \cos\theta \sin\phi d\theta d\phi \\
 &\quad (\text{symmetric}) + r^2 \sin^2\theta \cos^2\phi d\phi^2 + r^2 \sin^2\theta \cos^2\phi d\phi^2 \\
 dz &= \cos\theta dr - r \sin\theta d\theta \\
 dz^2 &= \cos^2\theta dr^2 - 2r \cos\theta \sin\theta dr d\theta + r^2 \sin^2\theta d\theta^2 \\
 ds^2 &= \sin^2\theta (\cos^2\phi + \sin^2\phi) dr^2 + 2r \sin\theta \cos\phi (\cos^2\theta + \sin^2\phi) dr d\theta \\
 &\quad + r^2 \cos^2\theta (\sin^2\phi + \cos^2\phi) d\theta^2 + r^2 \sin^2\theta (\sin^2\phi + \cos^2\phi) d\phi^2 + dz^2 \\
 &= \sin^2\theta dr^2 + 2r \sin\theta \cos\phi dr d\theta + r^2 \cos^2\theta d\theta^2 \\
 &\quad + r^2 \sin^2\theta d\phi^2 + \cos^2\theta dr^2 - 2r \cos\theta \sin\theta dr d\theta + r^2 \sin^2\theta d\theta^2 \\
 &= dr^2 + r^2 d\theta^2 + r^2 \sin^2\theta d\phi^2 \\
 &= dr^2 + r^2 (\sin^2\theta + \cos^2\theta) \quad \checkmark
 \end{aligned}$$

$$4.1 \quad E_1 = \frac{4}{3}\pi r^3 \epsilon_1 \quad \left\{ \begin{array}{l} \epsilon_1 = 5200 \text{ MeV m}^{-3} \\ r = 1 \text{ AU} = 1.5 \times 10^{11} \text{ m} \end{array} \right.$$

$$= \boxed{7.8 \times 10^{31} \text{ MeV}}$$

$$E_0 = \boxed{1.1 \times 10^{60} \text{ MeV}} \quad (\text{GOOGLE CALCULATOR})$$

$E_1 \ll E_0$ , should not have a significant effect.

4.2 If some matter is converted into radiation then the pressure of the universe increases since  $W_{\text{matter}} \ll W_{\text{radiation}}$  &  $\epsilon$  is constant.

By the acceleration equation this causes a to decrease w/ time, hence the universe contracts.

⇒ This is slightly unintuitive based on how we usually think about pressure.

$$4.3 \quad (4.69) \quad R_0 = \frac{c}{2\sqrt{\pi G P}} = (6.3 \times 10^{17} \text{ s}) c$$

↑

It would take a photon  $2\pi R_0 \sim 10^{18}$  sec to circumnavigate  $\sim 10^{11}$  years!

$$4.5 \quad E = \sqrt{m^2 c^4 + h^2 c^2 / \lambda^2} \quad E \text{ per one particle}$$

NONRELATIVISTIC :  $a \rightarrow 0 \Rightarrow \lambda \rightarrow \infty, E = mc^2$

$$\boxed{P = -\frac{\partial E}{\partial V} = 0}$$

RELATIVISTIC :  $a \rightarrow 0 \Rightarrow \lambda \rightarrow 0$   
 $\Rightarrow \frac{m^2 c^4}{h^2 c^2 / \lambda^2} \ll 1 \Rightarrow E \approx \frac{hc}{\lambda}$

$$\text{let } \lambda = \lambda_0 a = \lambda_0 V^{1/3}$$

$$E = \frac{hc}{\lambda_0} V^{-1/3}$$

$$P = -\frac{\partial E}{\partial V} = \frac{1}{3} \frac{hc}{\lambda_0} V^{-4/3} = \frac{1}{3} \frac{E}{V} = \boxed{\frac{1}{3} E}$$

ENERGY DENSITY

## EXTRA CREDIT

a) THE EASIEST WAY TO DO THIS IS TO PERFORM A LORENTZ TRANSFORMATION ON THE ENERGY:

$$E' = \gamma E - \beta \gamma c p_x$$

$$hv' = \gamma h\nu - \beta \gamma c (\gamma c \cos\theta)$$

$$v' = \gamma v(1 - \beta \cos\theta)$$

$$\boxed{\frac{v'}{v} = \frac{1 - \frac{v}{c} \cos\theta}{\sqrt{1 - \frac{v^2}{c^2}}}}$$

b) SET  $\frac{v'}{v} = 1$

$$1 - \frac{v}{c} \cos\theta = \sqrt{1 - \frac{v^2}{c^2}}$$

$$\boxed{\cos\theta = \frac{c}{v} (1 - \sqrt{1 - \frac{v^2}{c^2}})}$$