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Measurement of electron trapping in the Cornell Electron Storage Ring

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The buildup of low-energy electrons has been shown to affect the performance of a wide variety of particle accelerators. Of particular concern is the persistence of the cloud between beam bunch passages, which can impose limitations on the stability of operation at high beam current. We have obtained measurements of long-lived electron clouds trapped in the field of a quadrupole magnet in a positron storage ring, with lifetimes much longer than the revolution period. Based on modeling, we estimate that about 7% of the electrons in the cloud generated by a 20-bunch train of 5.3 GeV positrons with 16-ns spacing and 1.3×10^{11} population survive longer than $2.3 \mu\text{s}$ in a quadrupole field of gradient 7.4 T/m. We have observed a nonmonotonic dependence of the trapping effect on the bunch spacing. The effect of a witness bunch on the measured signal provides direct evidence for the existence of trapped electrons. The witness bunch is also observed to clear the cloud, demonstrating its effectiveness as a mitigation technique.

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I. INTRODUCTION

Electron cloud buildup has been observed in many accelerators since the 1960s [1]. Adverse consequences of electron cloud buildup include emittance growth, beam instabilities, and excess heat load to cryogenic systems.

Positron storage rings for which electron clouds have been an important factor in the design and performance include KEKB in Japan [2] and PEP-II in the USA [3]. Proton accelerators affected by electron clouds include the Los Alamos Proton Storage Ring (PSR) in the USA [4], CERN's Proton Synchrotron (PS), Super Proton Synchrotron (SPS) and Large Hadron Collider (LHC) [5]. At the LHC, electron cloud has been observed to affect the cryogenic heat load [6].

Electron cloud buildup is a major concern for accelerator upgrade programs and for the design of future accelerators. Electron cloud considerations have driven the design of the SuperKEKB collider [7] and the positron damping ring for the proposed International Linear Collider (ILC) [8]. The LHC luminosity upgrade is contingent on reducing the bunch spacing to 25 ns [9]; at this bunch spacing, severe electron cloud buildup has been observed, such that this bunch pattern has been used for beam scrubbing runs [5]. The success of the upgrade is likely to be contingent on limiting electron cloud buildup.

Considerable work has been done on the development of electron cloud mitigation techniques. At KEKB and PEP-II, solenoidal magnetic field windings were installed on the beam-pipes. For SuperKEKB, solenoidal windings are used in field-free regions, while TiN coatings and antechambers are included in quadrupole magnets, where solenoidal windings cannot be used. Carbon coatings for the dipole magnet vacuum chambers in the SPS are under study at CERN [10].

The electron cloud is observed to build up during the passage of a train of closely-spaced bunches, imposing restrictions on the operational bunch charge and train length. In field-free regions, gaps between trains allow the electron cloud to dissipate. In regions of magnetic field, however, cloud electrons can become trapped over long periods of time. Since trapped electrons can interact with the beam over many turns, they have the potential for more severe effects.

Electron cloud trapping has been studied experimentally and via simulation. Trapping of electrons oscillating around a 70-m-long proton bunch in the LANL PSR storage ring has been observed. [4]. At LBNL, electrons were observed to be trapped in the fields of an ion beam and accelerator elements, and measurements of the time dependence of electron cloud buildup were carried out [11]. Estimates of long-lived electron cloud buildup at the LHC and consequences for vacuum chamber heat load have been presented in Ref. [12]. More recently, heat load in the final-focus quadrupoles of the LHC has been attributed to electron cloud buildup [13]. Simulations were used to study electron trapping in quadrupole and sextupole magnets for the parameters of the KEKB positron ring [14], as well as for the Cornell Electron Storage Ring (CESR) and the ILC

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76 positron damping ring [15]. Prior to the measurements
 77 presented here, no experimental study of electron trapping
 78 in a positron storage ring has been available to validate
 79 modeling efforts.

80 A principal goal of the Cornell Electron Storage Ring
 81 Test Accelerator program [16] is to investigate performance
 82 limitations in future high-energy low-emittance rings.
 83 These studies include measurements of electron cloud
 84 buildup caused by synchrotron-radiation-induced photo-
 85 emission on the surface of the vacuum chamber. The CESR
 86 ring stores positron and electron beams of energy 1.8 GeV
 87 to 5.3 GeV, arranged in bunches spaced in intervals of 4
 88 or 14 ns, with bunch populations ranging up to 1.6×10^{11} .
 89 A variety of detectors sensitive to cloud electrons incident
 90 on the vacuum chamber wall have been used to study cloud
 91 buildup [17–21].

92 The potential for undesired consequences to accelerator
 93 performance motivated the study of electron trapping in
 94 the CESRTA electron cloud research program. In this paper,
 95 we report on the measurement of electron trapping in a
 96 quadrupole magnet over a $2.3 \mu\text{s}$ time interval between
 97 bunch train passages. Our demonstration of cloud trapping
 98 is based on two observations: first, the revolution-averaged
 99 electron flux arriving at the vacuum chamber wall during
 100 the passage of a ten-bunch train of positrons is greater when
 101 a second such bunch train immediately follows, showing
 102 that cloud is present at the time of arrival of the first ten-
 103 bunch train; second, inserting a single positron bunch over
 104 a broad time range centered halfway around the ring
 105 reduces the observed flux of electrons at the wall during
 106 the train passage, showing that trapped electrons were
 107 cleared by the intermediate bunch. It is noteworthy that
 108 beam-free intervals in the ring are ineffective at clearing the
 109 electrons, since the trapping mechanism is not contingent
 110 upon the beam potential as was the case at the PSR.

111 II. TIME-RESOLVING ELECTRON DETECTOR

112 Time-resolving electron detectors have provided detailed
 113 information on local cloud formation, allowing the inde-
 114 pendent characterization of photoelectron and secondary
 115 electron production mechanisms [20,21]. We have installed
 116 shielded detectors in a cylindrical stainless steel vacuum
 117 chamber of inner diameter 95.5 mm inside a 60-cm-long
 118 quadrupole magnet, as shown in Fig. 1a. One detector was
 119 located longitudinally near one end of the iron yoke in order
 120 to measure electron cloud buildup in the fringe field. In the
 121 following, we refer exclusively to measurements obtained
 122 from the detector positioned in the longitudinal center of the
 123 magnet and located in azimuth at 45 degrees from the
 124 horizontal mid-plane toward the inside of the ring, as shown
 125 in Fig. 1b. Electrons are collected on the 10-mm-wide copper
 126 trace (Fig. 1c) which tapers to a transmission line using the
 127 grounded copper on the other side of the 0.12-mm-thick
 128 Kapton sheet. The total length of the trace including the
 129 10-mm-wide, 102-mm-long rectangular central region is

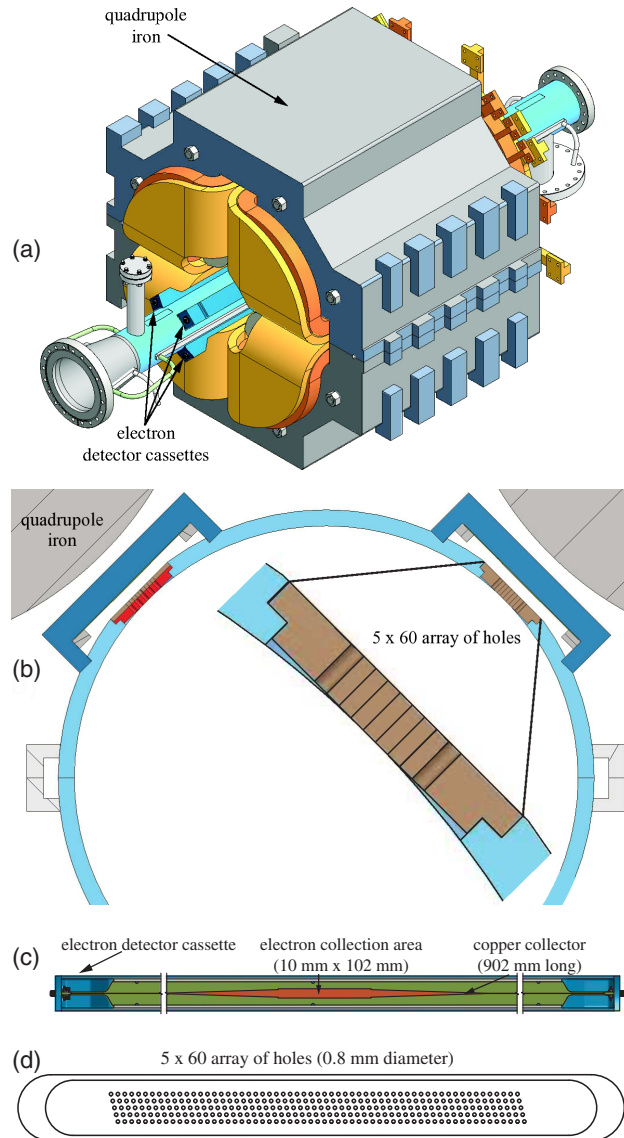


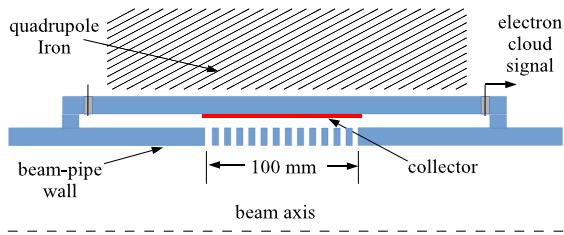
FIG. 1. (a) Vacuum chamber equipped with electron detectors
 in the quadrupole magnet. (b) Arrangement of two detectors
 in front of the magnet poles as seen from the positron arrival
 direction. (c) Geometry of the copper electrode biased at 50 V to
 collect electrons entering through the pattern of holes in the
 beam-pipe shown in (d). The rectangular region of the collector
 and the pattern of holes are each about 10 cm long.

907 mm. The pattern of 5×60 parallel 0.8-mm-diameter
 holes shown in Fig. 1d allows passage of cloud electrons
 through the beam-pipe to the collector. The chosen hole
 diameter gives a depth-to-diameter ratio of 3:1 in order to
 shield the detector from the rf power radiated by the 18-mm-
 long positron bunches [22]. The hole pattern is 7.1 mm wide
 and 94.4 mm long. Figure 2 shows a schematic view of the
 beam-pipe, hole pattern and detector arrangement.

The collector is biased at +50 V relative to the vacuum
 chamber in order to prevent secondary electrons from
 leaving the collector surface. The AC-coupled front-end

F1:1
 F1:2
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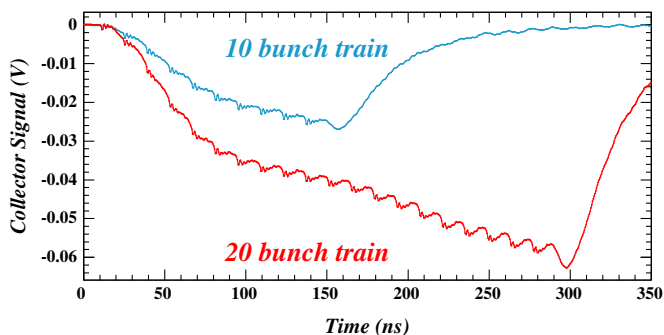


F2:1 FIG. 2. Schematic cross section of the electron detector, which
 F2:2 is located near the longitudinal center of the quadrupole magnet.
 F2:3 The holes in the beam-pipe wall allow cloud electrons to reach the
 F2:4 collector.

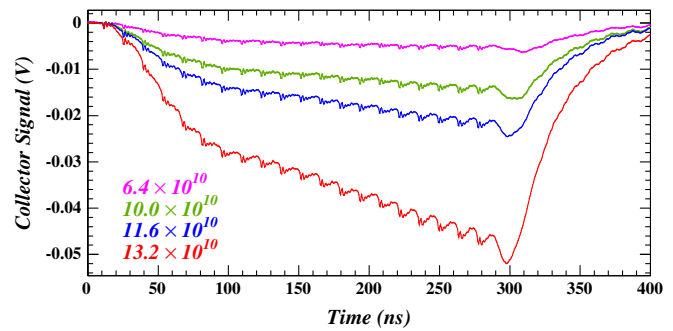
141 readout electronics consists of two Mini-Circuits ZFL-500
 142 broadband amplifiers with $50\ \Omega$ input impedance and a
 143 total gain of 40 dB. Oscilloscope traces are digitized to
 144 8-bit accuracy in 1000 time bins, typically 0.5 or 1.0 ns
 145 wide, averaging over 8000 beam-synchronous triggers. The
 146 direct beam-induced signal from the residual transmission
 147 of high-frequency rf power through the shielding holes
 148 results in a damped ringing in the raw oscilloscope signals.
 149 All signals depicted in the figures below show the result of
 150 applying a 13-MHz low-pass digital post-processing filter
 151 which suppresses this noise by an order of magnitude.

152 Figure 3 shows the filtered signals for 10- and 20-bunch
 153 trains of 5.3 GeV positrons. The bunches have rms sizes of
 154 1.8 mm horizontally and 0.08 mm vertically. The average
 155 bunch population is 1.3×10^{11} . The bunch spacing is 14 ns
 156 and the bunch-to-bunch population is uniform to a few
 157 percent. The quadrupole field gradient is 7.4 T/m, hori-
 158 zontally focusing.

159 The larger signal during the first 10 bunches of the
 160 20-bunch train relative to that for the 10-bunch train shows
 161 the presence of cloud prior to the arrival of the train. One
 162 can deduce that electrons remain trapped at least as long as
 163 the $2.3\ \mu\text{s}$ beam-free interval prior to the return of the
 164 bunch train. The decrease in cloud buildup rate following



F3:1 FIG. 3. Electron detector signals recorded for 10- and 20-bunch
 F3:2 trains of 5.3 GeV positrons for an average bunch population of
 F3:3 1.3×10^{11} . The enhanced signal during the first 10 bunches of a
 F3:4 20-bunch train relative to that for the 10-bunch train shows that
 F3:5 electrons were trapped during the entire $2.3\ \mu\text{s}$ interval prior to
 F3:6 the return of the bunch train.

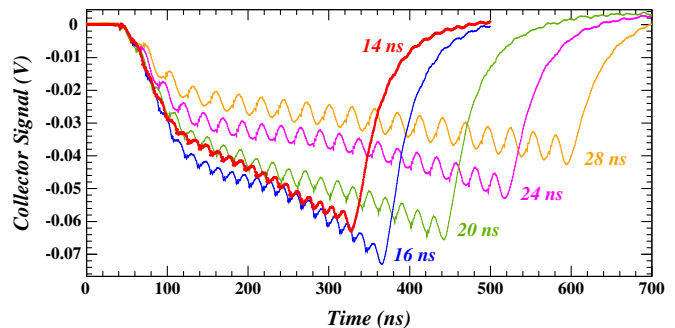


F4:1 FIG. 4. Dependence of the signals on bunch population for
 F4:2 20-bunch trains with 14 ns spacing. The dependence is strongly
 F4:3 nonlinear, the signal amplitude increasing by an order of
 F4:4 magnitude for a factor of two increase in bunch population.

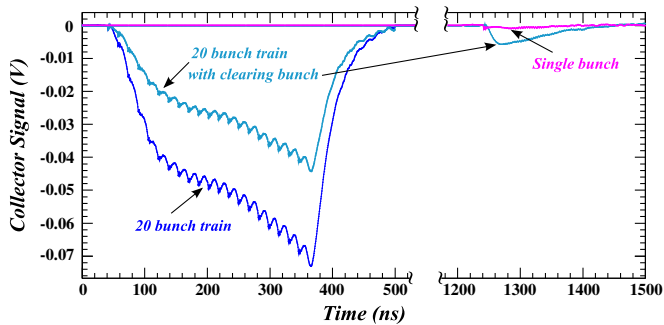
165 the first 6 bunches indicates that a subset of trapped
 166 electrons which can contribute signal has become depleted
 167 at that time. In spite of this clearing of the trapped reservoir
 168 of electrons, the signal does not return to the level of the
 169 10-bunch signal, showing that the additional cloud seeded by
 170 the long-term trapping is self-sustaining. The signal depends
 171 strongly on the bunch population, decreasing by an order
 172 of magnitude as the bunch population decreases by a factor
 173 of two from 1.3×10^{11} to 6.4×10^{10} , as shown in Fig. 4.

174 The dependence of trapping on the bunch spacing is
 175 shown in Fig. 5 for a bunch population of about 1.3×10^{11} .
 176 The decrease with increasing bunch spacing can be under-
 177 stood in terms of an overall decrease in cloud buildup.
 178 However, the enhancement of the signal at 16-ns spacing
 179 relative to the signal for 14-ns spacing shows that when
 180 electron trapping is of concern, care must be taken in the
 181 choice of bunch spacing.

182 We have investigated the effectiveness of an intermediate
 183 bunch as a mechanism for clearing the trapped cloud.
 184 Figure 6 shows the three signals obtained from (i) a
 185 20-bunch train, (ii) a 20-bunch train with a clearing bunch
 186 following about 900 ns after the end of the train, and (iii) a



F5:1 FIG. 5. Comparison of signals obtained from 20-bunch trains
 F5:2 with spacing 14, 16, 20, 24 and 28 ns. The increase in signal for
 F5:3 the 16-ns spacing relative to the 14-ns spacing shows that long-
 F5:4 term cloud electron trapping can be enhanced by an unfortunate
 F5:5 choice of bunch spacing.



F6:1 FIG. 6. Effect of an intermediate clearing bunch following
 F6:2 about 900 ns after the end of a 20-bunch train for the case of 16-ns
 F6:3 spacing. The difference in magnitude between the signals at
 F6:4 1250 ns is directly sensitive to the trapped electrons produced by
 F6:5 the 20-bunch train.

187 single bunch. The single-bunch signal is plotted to coincide
 188 with the signal from the clearing bunch for the purpose of
 189 comparison. The clearing bunch accelerates trapped cloud
 190 electrons into the detector, and thus provides direct evi-
 191 dence for the trapped cloud. In addition, the reduced signal
 192 from the 20-bunch train when the clearing bunch is present
 193 shows the effectiveness of such a mitigation technique. We
 194 verified that the clearing effectiveness is independent of the
 195 delay of the clearing bunch over a range of ± 500 ns.
 196 The full clearing effect was achieved when the clearing
 197 bunch population reached about 20% of the average
 198 population of the bunches in the train.

199 III. TRAPPING MECHANISM

200 The long-term trapping of electrons in nonuniform fields
 201 such as quadrupole fields can be understood in terms of an
 202 adiabatic magnetic moment μ given by

$$\mu = \frac{mv_{\perp}^2}{2B}, \quad (1)$$

203 where m is the mass of the electron, B is the magnetic field
 204 magnitude, and v_{\perp} is the velocity component perpendicular
 205 to the magnetic field vector (see, for example, Ref. [23]).
 206 This quantity remains invariant as long as $\frac{dB}{B} \ll 1$ during
 207 the cyclotron motion, or, equivalently,

$$\Gamma = \frac{|\nabla B|r_c}{B} \ll 1, \quad (2)$$

208 where r_c is the cyclotron radius. Combining the conditions
 209 of conservation of magnetic moment and conservation of
 210 energy, one can specify a “velocity-space loss cone” angle,
 211 Θ_{LC} , which defines the trapping condition. A particle
 212 moving from a region of lower field to a region of higher
 213 field reverses its path if the velocity components
 214 perpendicular and parallel to the magnetic field at the
 215 starting position, denoted by v_{\perp}^{in} and $v_{\parallel}^{\text{in}}$ respectively, are
 216 related such that

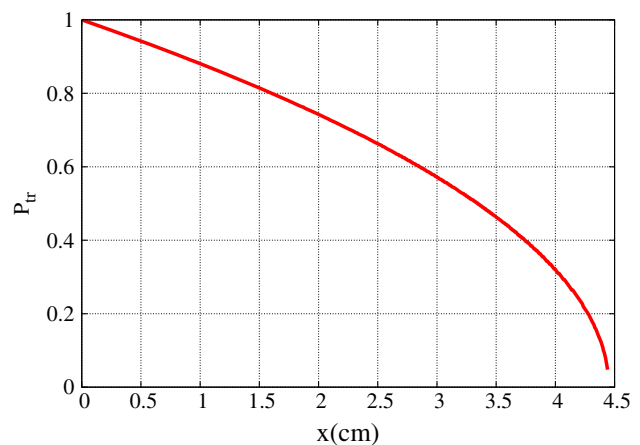
$$\frac{v_{\parallel}^{\text{in}}}{v_{\perp}^{\text{in}}} \leq \left(\frac{B_{\text{bd}}}{B_{\text{in}}} - 1 \right)^{1/2}. \quad (3)$$

217 Here B_{in} is the magnetic field magnitude at the start point,
 218 and B_{bd} is the magnitude along the field line at the
 219 boundary beyond which the particle is lost. If the above
 220 relationship is satisfied, the particle reaches a point where
 221 the parallel velocity goes to zero, and the particle reverses
 222 its path along the field line. In a quadrupole magnetic field,
 223 the trapped particle is confined between two such mirror
 224 points located along a field line symmetric about either the
 225 horizontal or the vertical axis. While the particle mirrors
 226 between the pair of points, it drifts in the longitudinal
 227 direction until it reaches the fringe region of the quadru-
 228 pole, where it can escape [24]. This drift is caused by a
 229 nonzero gradient and curvature in the magnetic field, often
 230 referred to as the “grad B” and “curvature” drift, respec-
 231 tively. For a 7.4 T/m field gradient, the longitudinal drift
 232 over the duration of one CESR beam revolution is
 233 significant only when the electron energy is of the order
 234 of 1 keV. The energy distribution obtained from the cloud
 235 build-up modeling described below indicate that less than
 236 3% of the electrons have energies exceeding 1 keV.

237 The cosine of the loss cone angle represents the frac-
 238 tional solid angle in velocity space within which a particle
 239 remains confined. Thus, for a localized distribution of
 240 isotropic velocities, it represents the probability of confine-
 241 ment at that point. It can be expressed as

$$P_{\text{tr}} = \cos \Theta_{LC} = \left(1 - \frac{B_{\text{in}}}{B_{\text{bd}}} \right)^{1/2} \quad (4)$$

242 and is shown in Fig. 7 as a function of horizontal position x
 243 along the midplane of the vacuum chamber. The probability
 244 of confinement decreases with x , the distance from the
 245 beam, provided $\Gamma \ll 1$. The adiabatic condition can be
 246 expressed as



F7:1 FIG. 7. Cosine of the loss-cone angle, P_{tr} , versus horizontal
 F7:2 position in the mid-plane of the vacuum chamber. The trapping
 F7:3 probability increases toward the center of the chamber as long as
 F7:4 the adiabaticity condition is satisfied.

$$\Gamma = \frac{\sqrt{2mE_{\perp}}/e}{Kx^2} \ll 1, \quad (5)$$

247 where e is the electron charge, K is the quadrupole field
 248 gradient and E_{\perp} the kinetic energy corresponding to the
 249 velocity component perpendicular to the magnetic field.
 250 For electrons in a quadrupole with field gradient
 251 $K = 7.4$ T/m, Γ reduces to

$$\Gamma = 4.6 \times 10^{-3} \frac{\sqrt{E_{\perp}/\text{eV}}}{(x/\text{cm})^2}. \quad (6)$$

252 For comparison, the beam kick produced by a bunch
 253 carrying 1.3×10^{11} positrons on an electron at the vacuum
 254 chamber wall is 60 eV in the impulse approximation [25],
 255 easily satisfying the trapping condition. On the other hand,
 256 an electron with a horizontal momentum of 40 keV/c
 257 located 1 cm from the beam in the horizontal midplane is
 258 likely to hit the chamber wall.

259 IV. NUMERICAL MODELING OF ELECTRON 260 CLOUD BUILDUP

261 We have employed a particle-in-cell, time-sliced cloud
 262 buildup modeling code [26] to improve our understanding
 263 of the electron trapping mechanism and the observed
 264 signals. The code includes simulation algorithms for
 265 photoelectron generation, macroparticle tracking in the
 266 2D electrostatic fields of the beam and the cloud, and
 267 3D tracking in a variety of ambient magnetic fields, as well
 268 as for a detailed model of the interaction of cloud electrons
 269 with the vacuum chamber surface [27].

270 The code has been supplemented with response func-
 271 tions for the time-resolving electron detectors [28]. As a
 272 function of incident angle and energy, a fraction of the
 273 macroparticle charge hitting the wall in the region of the
 274 detector contributes to the modeled signal. The fraction is
 275 derived from an analytic calculation of the hole acceptance
 276 for the case of a magnetic field parallel to the hole axis. For
 277 an arbitrary magnetic field strength, the acceptance of the
 278 holes is derived by relating the incident kinetic energy and
 279 angle to the cyclotron radius and the wall traversal time,
 280 i.e., the fractional number of cyclotron revolutions per-
 281 formed in the wall. Thus the acceptance at high field
 282 extends to grazing angles of incidence when the cyclotron
 283 radius is smaller than the hole radius.

284 The amplitude of the modeled signal was found to be
 285 very sensitive to the assumed secondary emission yield,
 286 increasing by an order of magnitude as the peak secondary
 287 yield was increased from 1.4 to 1.9. The measured signal
 288 amplitude was reproduced with values for the peak sec-
 289 ondary yield and elastic yield of 1.4 and 0.5, respectively.

290 The model shows the signal to be generated predomi-
 291 nantly by electrons originally produced on the field lines
 292 entering the detector, i.e., from a narrow surface region in

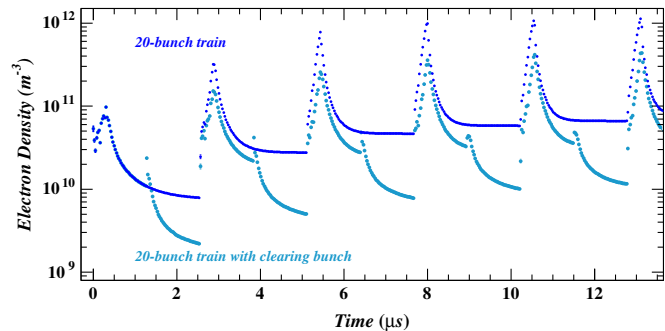


FIG. 8. Results for the beam-pipe averaged cloud density from the numerical model of electron buildup for the case of the 5.3 GeV 20-bunch train of positrons with 16-ns spacing shown in Fig. 6, with and without an intermediate clearing bunch.

front of the diametrically opposed pole and from 4-mm-wide regions on the vacuum chamber surface in front of the other two poles extending from the middle of the pole toward the detector. These signal macroparticles spiral around field lines which pass within a few millimeters of the beam. The electrons which remain trapped during the 2.3 μs prior to the train arrival are cleared out during the first 6 of the 20 bunch passages, reabsorbed either in the detector or the vacuum chamber wall. The signal also shows that the cloud development proceeds at the higher density level following the clearing, since it does not return to the level of the signal for a 10-bunch train. The trapping results in a sustained higher cloud density even after the trapped electrons have been removed.

Figure 8 shows the modeled electron cloud density averaged over the test volume of the cylindrical vacuum chamber for the case of a 20-bunch train of positrons with average bunch population 1.3×10^{11} , with and without an intermediate clearing bunch of the same population.

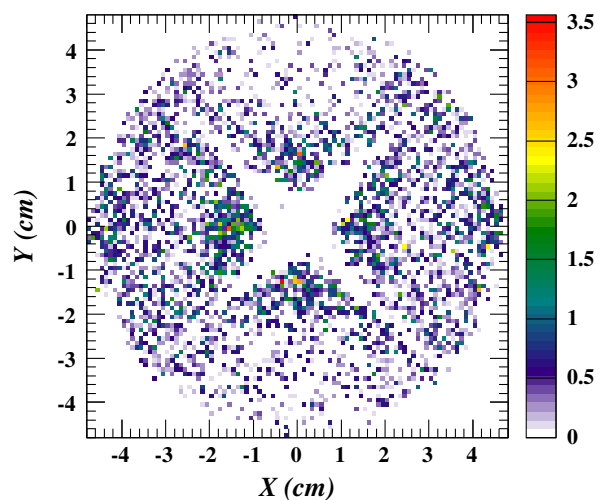


FIG. 9. The modeled transverse distribution of the trapped cloud shown at the end of the first beam revolution. The color scale ranges up to a maximum of 3.5×10^6 electrons/bin.

312 The peak density in the absence of the clearing bunch
 313 reaches $1.1 \times 10^{12} \text{ m}^{-3}$ after three turns, about 7% of
 314 which is trapped until the train returns. The clearing
 315 bunch reduces the trapped cloud density by about a factor
 316 of four.

317 The modeled transverse distribution of the cloud trapped
 318 in the quadrupole magnet at a time immediately preceding
 319 the return of the train is shown in Fig. 9. The trapped
 320 electrons are concentrated in four quadrants near the beam
 321 outside of a central depletion zone of 2 cm diameter,
 322 consistent with the trapping probability distribution shown
 323 in Fig. 7 and the nonadiabaticity in the central and diagonal
 324 regions. The median energy of the trapped electrons is
 325 about 50 eV.

326 V. SUMMARY

327 Our measurements with a time-resolving electron
 328 detector located in a quadrupole magnetic field have
 329 provided comparisons of signals from 10- and 20-bunch
 330 trains of positrons which show clear evidence for electron
 331 trapping during the entire $2.3 \mu\text{s}$ time interval prior to the
 332 return of the bunch train. Modeling tuned to the recorded
 333 signals indicates that approximately 7% of the cloud
 334 generated by a 5.3 GeV train of 20 bunches, each carrying
 335 1.3×10^{11} positrons, remains trapped. The measurements
 336 show a nonmonotonic dependence on bunch spacing. The
 337 clearing effect of an intermediate bunch has been mea-
 338 sured and successfully modeled, showing the trapped
 339 cloud can be reduced by a factor of four by such a
 340 clearing bunch. This characteristic of a quadrupole mag-
 341 netic field to concentrate electrons near the beam raises
 342 concerns for storage rings with positively charged beams,
 343 since those electrons can be attracted into the beam. Such
 344 measurements quantifying electron trapping in quadru-
 345 pole magnets provide information useful for the develop-
 346 ment of simulation codes which serve to predict electron
 347 cloud phenomena in future accelerators and to aid in the
 348 design of mitigation techniques.

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