

## An introduction to AdS/CFT and Klebanov-Strassler

8

### I. Warping and the holographic principle

Let's consider a Schwarzschild black hole in 4d. It carries a mass  $M$  and has - in the classical limit - an event horizon of radius  $R \propto M$ . One can argue (Bekenstein) that the entropy of the BH is

$$S = \frac{A}{4G_N k}$$

[1]

Roughly, the entropy is a measure of the information content.

One might have naively expected the entropy to grow with the volume and not the area.

This is the basis of the holographic principle ('t Hooft, Susskind) which roughly states that ~~the~~ a gravitational theory in a spacetime region  $\mathcal{M}$  is encapsulated by ~~physics~~ a theory

on the boundary  $\partial\mathcal{M}$ .

The most precise implementation

of this is the gauge/gravity

correspondence. The correspondence consists of pairs of the type

gravity on  $\mathbb{R}^{d-1} \times_w X^q \leftrightarrow$  gauge theory on  $\mathbb{R}^{d-1}$

[2]

Clarification of " $\leftrightarrow$ " is one of the goals of these lectures.

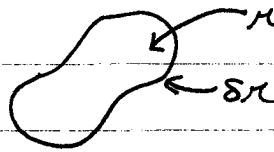
$\times_w$  denotes a warped product. It means that the spacetime isn't a direct product but that the metric takes the form (mostly positive signature)

$$ds_p^2 = e^{2A(y)} \eta_{\mu\nu} dx^\mu dx^\nu + e^{2\phi A(y)} g_{mn}(y) dy^m dy^n$$

[3]

$g_{mn}(y)$  is a metric on  $X^q$  and  $\eta_{\mu\nu}$  is the flat  $d$ -dimensional Minkowski

direction.  $\phi$  is some convenient constant.  $A(y)$  (or  $e^{nA}$ ) is called the warp factor.



A particularly important example is anti-de Sitter space (AdS).  $\text{AdS}^{d+1}$

$$\text{AdS}^{d+1} = R^{d-1} \times_{\text{hyp}} R^{>0}$$

$$ds_{d+1}^2 = \left(\frac{r}{L}\right)^2 dx_a^2 + \left(\frac{r}{L}\right)^2 dr^2$$

[4]

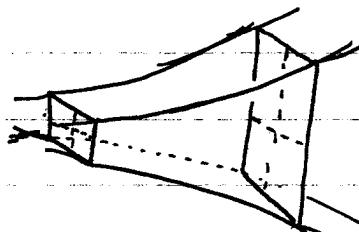
$y = e^{-r/L}$  and  $z = \frac{r}{L}$  are often

used. AdS has many interesting properties, few of which

we will use here, but

are important for a

deeper understanding of the duality.



AdS plays a central role in the correspondence.

Example:

~~AdS~~ type IIB supergravity

$N=4 \text{ SU}(N)$

[5]

$$\text{on } \text{AdS}^5 \times S^5 \cong R^{3,1} \times_{\text{hyp}} R^6$$

SYM

$N=4 \text{ SU}(N)$  SYM is an example of a conformal theory (i.e. one invariant under all angle-preserving transformations including coordinate rescaling). More generally,

$$\text{gravity on } \text{AdS}^d \times X \xleftrightarrow{\text{CFT}} \text{CFT on } R^{d-1}$$

[6]

These examples go under the name AdS/CFT. The term is often used as a metonym for the ~~&~~ Correspondence as a whole.

My personal bias is toward phenomenological application, so I will focus on  $d=4$  examples. There is rich physics to be found with other  $d$ .

If sufficiently limited questions are asked, we do not need ~~string theory~~ for quantum gravity. However, I will consider stringy examples because

1. The duality is most clearly "derived" from strings.

2. The duality cannot possibly be exact with classical gravity

3. Personal bias.

## II. Elements of string theory

For  $d=4$ , type-IIB string theory works best (though Sakai-Sugimoto, an attempt at AdS/QCD uses IIA). String theory isn't a spacetime field theory, but at low energies and weak coupling it has an effective field theory description.

For type-IIB string theory, the eft is type-IIB supergravity.

IIB sugra is the  $N_{10} = 10,2$  theory. Its field content is completely fixed by supersymmetry:

|                     |                        |  |                         |
|---------------------|------------------------|--|-------------------------|
| metric              | $\hat{g}_{\mu\nu}$     | symmetric tensor                           | Neveu-Schwarz<br>sector |
| Rabi-Ramond field   | $B_{\mu\nu}$           | anti-symmetric tensor                      |                         |
| dilaton             | $\phi$                 | scalar                                     |                         |
| axion               | $C$                    | scalar                                     | Ramond-Ramond sector    |
| RR 2-form potential | $C_{\mu\nu}$           | anti-symmetric tensor                      |                         |
| RR 4-form potential | $C_{\mu\nu\rho\sigma}$ | anti-symmetric tensor                      |                         |
| gravitini           | $\psi_m^{i=1,2}$       | left-handed Majorana-Weyl-Rarita-Schwinger | NS-R sector             |
| dilatini            | $x^{i=1,2}$            | left-handed Majorana-Weyl                  |                         |

These modes come from the lightest excitations of closed strings. ST has an infinite tower of modes with mass splittings set by the string tension

$$\propto \tau_{FI} \sim \alpha' \quad \Delta m^2 \sim \frac{1}{\alpha'} \sim \frac{1}{L^2}$$

[7]

The coupling strength of strings is set dynamically

$$g_{\text{RS}} = g_s e^\phi$$

[8]

(NB: notational differences differ greatly here)

To get from ST to sugra, we take the limits in the following way

Quantum string theory

$$\downarrow g_{\mu\nu} \rightarrow \infty \quad g_{\alpha\beta} \rightarrow 0$$

Classical string theory

$$\downarrow p^2 \alpha' \rightarrow 0$$

Classical supergravity

Note: M-theory and F-theory gives us a way to take  $p^2 \alpha' \rightarrow 0$  before  $g_{\alpha\beta} \rightarrow 0$ .

### II.A. Black p-branes

Consider Einstein-Maxwell theory in 4d. The equations of motion admit solutions for massive charged black holes (Reissner-Nordström). They admit a particularly limit called the extremal limit in which  $M=Q$  in some appropriate units. These branes are then stable - Hawking radiation will no longer occur.

The type-II theories have analogues to the Maxwell fields, namely the RR and NS potentials.

A black p-brane is analogous to a charged black-hole; a black D<sub>p</sub>-brane will couple electrically to  $C_{(p+1)}$  for  $p \leq 3$  and magnetically to  $C_{(7-p)}$  for  $p \geq 3$ . In IIB, we hence have stable ~~branes~~ black p-branes for odd p (In IIA we have even p). ~~These electric charges are at integer values~~

The black 3-brane solution is (Townsend; Horowitz, Strominger)

$$ds_{10}^2 = e^{2A} \eta_{\mu\nu} dx^\mu dx^\nu + e^{-2A} (dr^2 + r^2 ds_3^2)$$

$$\not \rightarrow e^A = 1$$

[9]

$$C_{\mu\nu\rho\sigma} = g_3^{-1} e^{4A} \epsilon_{\mu\nu\rho\sigma}$$

$$e^{-4A} = 1 + \frac{L^4}{r^4} \quad L^4 = 4\pi g_s N \alpha'^2$$

Note that a D3 couples both electrically and magnetically to  $C_4$ , so

there are additional components of  $C_4$  that cannot be globally defined

(c.f. - Dirac string for a magnetic monopole in 4d).

$N$  is the amount of D3 charge. Recall in the Maxwell theory that the amount of charge in a region is

$$Q \propto \int_{\Sigma} \vec{E} \cdot d\vec{A} \sim \int_{\Sigma} F_{ij} dx^i dx^j \sim \int F_{00}$$

[10]

Analogously, we define

$$F_{MNPRS} \tilde{\equiv} \partial_{[M} C_{NP]S}$$

[11]

Then,

$$N \stackrel{\text{def}}{=} \frac{1}{4\pi^2 \alpha'^2} \int F_{00}$$

[12]

Black p-branes can be thought of as solitonic configurations of closed strings.

It is possible to consider finite temperature black branes. These are very relevant for applications to condensed-matter and heavy ion physics, but we will not use them here.

Black p-branes are "half-BPS." This means many things, but here I mean

IIIBR<sub>32</sub>-branes that it preserves half of the supersymmetries. That is, IIB has 32 supercharges, but only half of them annihilate the state corresponding to a black Dp-brane.

## II.B. - D<sub>p</sub>-branes

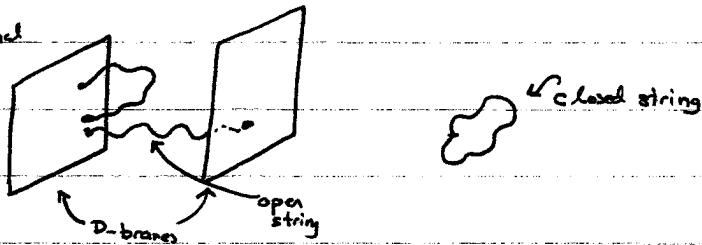
In addition to closed strings, type-II theories contain open strings.

These end on (p+1)-dimensional

surfaces called D<sub>p</sub>-branes

(the D is for Dirichlet)

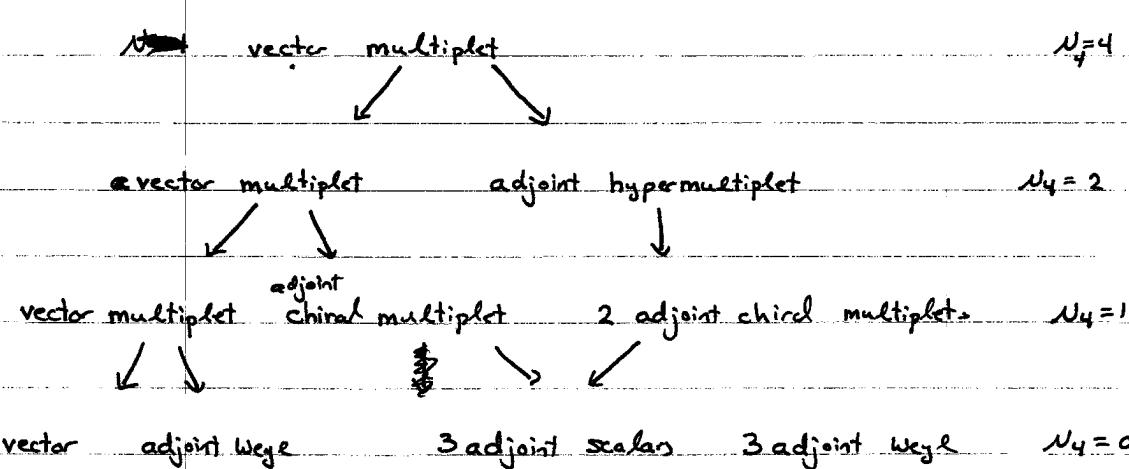
(Dai, Leigh, Polchinski).



The lightest modes of an open string on a coincident stack of N D<sub>p</sub>-branes

in flat space give a (p+1)-dimensional U(N) SYM theory with 16 supercharges.

For p=3, this is N=4 SYM.



In N=1 language, the superpotential is

$$W = g_{YM} \epsilon_{ijk} \bar{\Phi}^i \bar{\Phi}^j \bar{\Phi}^k$$

[13]

There are also many massive open-string states.

The coupling constant for a D3 is  $\sim g_s e^{\frac{2\pi}{\alpha}} \sim g_{YM}^2$ .

Let's recall that  $g_{YM}$  is not the correct expansion parameter for QCD - it's  $\Lambda$ .

Instead, loops involve summing over colors and so the correct expansion is the 't Hooft parameter

$$\lambda = g_{YM}^2 N^2$$

[14]

The scalars on the D3-world volume correspond to transverse fluctuation of the Dp-brane. That is, a Dp-brane is a dynamical object and the dynamics are described by open string theory.

A non-trivial fact is that Dp-branes carry Ramond-Ramond charge (Polchinski).

~~The most sensible interpretation is that~~ In fact the charge and tension match those of a black p-brane. The most sensible ~~an~~ interpretation is that p-branes and black p-branes are the same thing.

String theory is not a theory of strings alone, but a theory of strings and branes. Indeed there are a number of dualities between string theories that map strings and branes to each other so that there is no invariant distinction between them.

### III. Basics of AdS/CFT

Let's return to the black 3-brane picture.

$$ds_{10}^2 = \left(1 + \frac{r^2}{L^2}\right)^{-1} dx_4^2 + \left(1 + \frac{r^2}{L^2}\right) (dr^2 + r^2 ds_5^2)$$

[15]

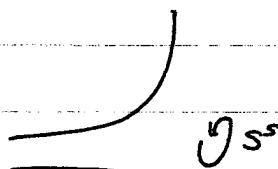
Note that in writing this picture we have implicitly fixed the position of the 3-branes. There are two interesting regions.

When  $r \gg L$ , the space

is approximately flat  $\mathbb{R}^9$ !

When  $r \ll L$ , it is strongly

warped  $AdS^5 \times S^5$



$$e^{-4A} \rightarrow \frac{1}{r^4} \quad r \gg L$$

consistency check

One can show (Maldacena) that in the low-energy limit these two sectors decouple. Moreover since we have fixed the positions of the 3-branes, there are no intermediate modes. Hence in the black-brane picture we have that theory is

$$\text{IIB sugra on } \text{AdS}^5 \times S^5 \otimes \text{IIB sugra on } R^{9,1}$$

[16]

Now let's consider

Now let's consider the open-string picture. The gauge theory is

$$U(N) \approx SU(N) \times U(1)$$

[17]

The  $U(1)$  corresponds to fluctuations of the center-of-mass and so by fixing the position of the D3s we essentially just ignore the  $U(1)$ . The theory consists of open strings on the D3s and closed strings in  $R^{9,1}$ . In the low energy limit, there are no interactions between them and so the theory is

$$N=4 \text{ } SU(N) \text{ gauged IIB } \otimes \text{IIB sugra on } R^{9,1}$$

[18]

Since black ~~for~~ 3-branes are D3-branes we must have at low energies,

$$\boxed{\text{IIB sugra on } \text{AdS}^5 \times S^5 \Leftrightarrow N=4 \text{ } SU(N) \text{ SYM.}}$$

[19]

Now let's make a couple of observations:

1.  $N=4 \text{ } SU(N) \text{ SYM}$  is a CFT so there is no meaning to "low the phrase  
"low energy"

2.  $\text{AdS}^5 \times S^5$  is a valid solution to IIB even without ever having to make the decoupling argument.

We thus <sup>conclude</sup> ~~suppose~~ that [19] must hold at all energies.

10.

The duality is a weak-strong duality in the sense that when  $\lambda \gg 1$ , the gravity side is weakly curved and so  $\alpha'$  corrections are small. On the other hand, when  $\lambda \ll 1$ , the gauge theory is weakly coupled but  $\alpha'$  effects are important.

11.

Here for the duality to always hold, ST must come in. We will later show that even for  $\lambda \gg 1$ , the duality is not complete without ST.

Here's what we've learned so far:

IIB ST on  $AdS^5 \times S^5 \Leftrightarrow N=4$  SW(N) SYM

$$\alpha' \frac{1}{L^2} \Leftrightarrow \lambda_S$$

$$g_S e^{\phi} \Leftrightarrow g_M$$

### III. A - Matching symmetries

An important check of the duality is the matching of symmetries. A conformal group field theory is invariant under

1. Lorentz transformations  $x^\mu \mapsto x^\mu + \omega^\mu_\nu x^\nu$

2. Translations  $x^\mu \mapsto x^\mu + a^\mu$

3. Dilations  $x^\mu \mapsto x^\mu + 2x^\mu$

4. Special transformations  $x^\mu \mapsto x^\mu + x^\nu b^\mu_\nu - 2b^\mu x^\mu$

The resulting group is  $SO(4,2)$ . (Note: In  $d=2$  the group is much larger.)

This matches onto the  $SO(4,2)$  isometry of  $AdS^5$ . The ~~SO(5)~~ Poincaré subgroup corresponds to that of the CFT. We have also

$$(x^\mu, r) \mapsto (\lambda x^\mu, \lambda^{-1} r)$$

[26]