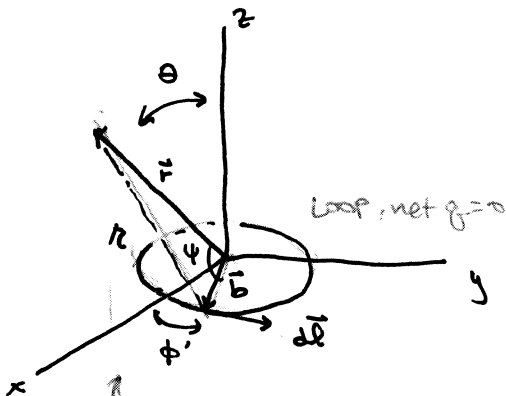


AGENDA

- ANNOUNCEMENTS: HW #3 / MIDTERM GRADING ASAP, YUEN
- MAGNETIC DIPOLE RADIATION (brief)
- POINT CHARGES
- RELATIVITY PRIMER
- PROBLEMS

MAGNETIC DIPOLE RADIATION



$$I(t) = I_0 \cos(\omega t) \quad \rightarrow \quad = \pi b^2 I_0$$

$$\vec{m}(t) = \pi b^2 I(t) \hat{z} = m_0 \cos(\omega t) \hat{z}$$

MAGNETIC DIPOLE $\frac{\omega b r}{c}$

$$\vec{A}(\vec{r}, t) = \frac{\mu_0}{4\pi} \int \frac{I_0 \cos(\omega(t - r/c))}{r} d\vec{l}$$

SUMMERIZE

$$\vec{A}(\vec{r}, t) = \frac{\mu_0 I_0 b}{4\pi} \int_0^{2\pi} \frac{\cos(\omega(t - r/c))}{r} d\phi'$$

By symmetry can use
 $\vec{r} = r \sin\theta \hat{x} + r \cos\theta \hat{z}$
 $b = b \cos\phi' \hat{x} + b \sin\phi' \hat{y}$

① → DETERMINE APPROXIMATIONS

② DETERMINE \vec{A} USING

$$\vec{A}(\vec{r}, t) = \frac{\mu_0}{4\pi} \int \frac{\vec{J}(\vec{r}', t')}{r} d\tau'$$

- $b \ll r$
- $b \ll \lambda$
- $r \gg \lambda$

$$\lambda = 2\pi \frac{c}{\omega} \quad (\text{dimless})$$

① DETERMINE r USING LAW OF COSINES

$$r^2 = r^2 + b^2 - 2rb \cos\phi'$$

$$= 2\vec{r} \cdot \vec{b} = r b \sin\theta \cos\phi'$$

② LOOP SMALLER THAN OBS. DISTANCE ($\frac{b}{r}$ is small)

$$r^2 \approx r^2 (1 - \frac{2b}{r} \sin\theta \cos\phi') \quad \leftarrow \text{PROP } o(\frac{b^2}{r^2})$$

$$r \approx r \sqrt{1 - \epsilon} \approx r (1 - \frac{1}{2}\epsilon) \quad \leftarrow \epsilon = \frac{2b}{r} \sin\theta \cos\phi'$$

$$\frac{1}{r} = \frac{1}{r} \cdot \frac{1}{1 - \epsilon} \approx \frac{1}{r} (1 + \epsilon')$$

$$\frac{1}{r} = \frac{1}{r} \left(1 + \frac{b}{r} \sin\theta \cos\phi' \right)$$

$$= \frac{1}{r} \left(1 + \frac{1}{2}\epsilon \right)$$

2a) ALSO APPROXIMATE $\cos(\omega t_r)$

$$\cos(\omega(t - r/c)) \approx \cos\left[\omega t - \frac{1}{c} \left\{ r \left(1 - \frac{1}{2}\epsilon \right) \right\}\right]$$

$$\cos\left(\omega\left(t - \frac{r}{c}\right) + \frac{\omega}{2c}\epsilon\right)$$

CALL THIS $\tilde{\omega} t_r$

2b) USE $\cos(a+b) = \cos a \cos b - \sin a \sin b$

$$\cos(\omega t_r) \approx \underbrace{\cos(\omega(t - r/c))}_{\tilde{\omega} t_r} \cos\left[\frac{\omega\epsilon}{2c}\right] - \underbrace{\sin(\omega(t - r/c))}_{\tilde{\omega} t_r} \sin\left[\frac{\omega\epsilon}{2c}\right]$$

3) SIZE OF DIPOLE SMALL VC. WAVELENGTH

NOTE: $\epsilon = \frac{2b}{r} \sin\theta \cos\phi'$

SO $\frac{\omega\epsilon}{2c} \propto b/(c/w)$

$\cos \omega\epsilon/2c \rightarrow 1$
 $\sin \omega\epsilon/2c \rightarrow \omega\epsilon/2c$

$$\cos(\omega t_r) \approx \underbrace{\cos(\omega(t - r/c))}_{\tilde{\omega} t_r} - \frac{\omega b}{c} \sin\theta \cos\phi' \underbrace{\sin(\omega(t - r/c))}_{\tilde{\omega} t_r}$$

4) DUG INTO $\vec{A}(r, t) = k \hat{y} \int \cos(\omega t_r)/r \cos\phi' d\phi'$

$$= k \hat{y} \int \left\{ \underbrace{\cos(\omega(t - r/c))}_{\tilde{\omega} t_r} - \frac{\omega b}{c} \sin\theta \underbrace{\sin(\omega(t - r/c))}_{\tilde{\omega} t_r} \right\} \frac{1}{r} \left(1 + \frac{1}{2}\epsilon \right) \cos\phi' d\phi'$$

$$\approx \frac{k}{r} \hat{y} \int \left\{ \underbrace{\cos(\omega \tilde{t}_r)}_{\text{INTEGR. TO ZERO}} + \frac{1}{2}\epsilon \cos(\omega \tilde{t}_r) - \frac{\omega b}{c} \sin\theta \underbrace{\sin(\omega \tilde{t}_r)}_{\text{HAS } \phi' \text{ dependence}} \right\} \cos\phi' d\phi'$$

$$\approx \frac{k}{r} \hat{y} \left(\frac{b}{r} \sin\theta \right) \int \cos^2\phi' \left(\cos(\omega \tilde{t}_r) - \frac{\omega b}{c} \sin\theta \sin(\omega \tilde{t}_r) \right) d\phi'$$

↑
INTEGRATES TO π

⇒ turn $\hat{y} \rightarrow \hat{\phi}$

↳ $\vec{A} = \frac{\pi k b^2}{r^2} \sin\theta \left(\cos(\omega \tilde{t}_r) - \frac{\omega b}{c} \sin\theta \sin(\omega \tilde{t}_r) \right) \hat{\phi}$

5) $r \gg c/\omega$ IN RADIATION ZONE

so $\frac{r}{c} \ll 1 \Rightarrow \cos(\omega t - r/c) \ll \frac{r}{c} \sin(\omega t - r/c)$

$$\Rightarrow A(r, \theta, t) = -\frac{E_0}{r^2} \sin \theta \frac{\omega r}{c} \sin(\omega t - r/c) \hat{\phi}$$

$$= \boxed{-\frac{\mu_0 I_0 b^2}{4\pi c} \frac{\sin \theta}{r} \sin(\omega t - r/c) \hat{\phi}}$$

find :

$$\vec{E} = \frac{\mu_0 m_0 \omega^2}{4\pi c} \left(\frac{\sin \theta}{r}\right) \cos(\omega t - r/c) \hat{\phi}$$

$$\vec{B} = -\frac{\mu_0 m_0 \omega^2}{4\pi c^2} \left(\frac{\sin \theta}{r}\right) \cos(\omega t - r/c) \hat{\theta}$$

ALWAYS TRUE } \rightarrow IN PHASE, \perp , TRANSVERSE TO PROPAGATION \hat{r}

$\rightarrow E_0/B_0 = c$

\rightarrow LOOK JUST LIKE E. DIPOLE w/ $m_0 \leftrightarrow cp_0$
 $(\hat{\theta}, \hat{\phi}) \leftrightarrow (\hat{\phi}, -\hat{\theta})$

$$\langle \vec{S} \rangle = \frac{\mu_0 m_0^2 \omega^4}{32\pi^2 c^3} \frac{\sin^2 \theta}{r^2} \hat{r}$$

$$\langle P \rangle = \frac{\mu_0 m_0^2 \omega^4}{12\pi c^3}$$

$$P_M/P_E = \left(\frac{m_0}{P_0 c}\right)^2$$

ARBITRARY SOURCE RADIATION

- THIS MAY BE THE PINNACLE OF USUAL PHYSICS IN THE BOOK
- YOU SHOULD COVER THIS, IT WOULD TAKE US FOREVER !!
- NOTE TO NOTE: USE OF CONS. OF CHARGE
→ READ HIS COMMENTS ON P. 418 FIRST

IDEA: MULTIPOLE EXPANSION

IF ELEC. 2^N -POLE VANISHES (OR $\frac{d^2}{dt^2}$ IS ZERO)
MUST USE @ NEXT TERM, (LHM) @ IN r'
→ WHICH IS RELATED TO $\left\{ \begin{array}{l} \text{MAGNETIC } 2^N\text{-POLE} \\ \text{ELECTRIC } 2^{(N+1)}\text{-POLE} \end{array} \right.$

POINT CHARGES

From last time:

$$\vec{E}(\vec{r}, t) = \frac{q}{4\pi\epsilon_0} \frac{1}{(\vec{r} \cdot \vec{u})^3} \left[(c^2 - v^2) \vec{u} + \vec{r} \times (\vec{u} \times \vec{a}) \right]$$

$$\vec{u} = c\hat{r} - \vec{v}$$

$$\vec{E} = \frac{1}{r} \hat{r} \cdot \vec{E}(\vec{r}, t)$$

$$\vec{S} = \frac{1}{\mu_0 c} [E^2 \hat{r} - (\hat{r} \cdot \vec{E}) \vec{E}]$$

terrible analogy: college students

RECALL ARGUMENT:

RADIATION GOES TO ∞ (flies that leave garbage truck)
SURFACE AREA $\propto r^2$
SO E. FLUX MUST GO AS $1/r^2$ OR MORE SLOWLY
UNPHYSICAL

$\vec{S} \propto E^2 \Rightarrow$ PART OF \vec{E} THAT GOES AS $1/r$
IS RADIATION

$$\vec{E}_{rad} = \frac{q}{4\pi\epsilon_0} \frac{1}{(\vec{r} \cdot \vec{u})^3} \left[\vec{r} \times (\vec{u} \times \vec{a}) \right]$$

FOR WAVE @ REST @ $t = t_r$, $\vec{u} = c\hat{r}$

$$\vec{E}_{rad} = \frac{M \cdot q}{4\pi\epsilon_0 r} \left[(\hat{r} \cdot \vec{a}) \hat{r} - \vec{a} \right] + \frac{q}{4\pi\epsilon_0 c r} \left[\hat{r} \times (\hat{r} \times \vec{a}) \right]$$

$$\Rightarrow \vec{S}_{rad} = \frac{M \cdot q^2 a^2}{16\pi^2 \epsilon_0 c} \left(\frac{4\pi^2 \epsilon_0}{\mu_0} \right) \hat{r} = \frac{1}{4\pi\epsilon_0} \frac{q^2 a^2}{4c^3 r^2} \left[a^2 - (\hat{r} \cdot \vec{a})^2 \right] \hat{r}$$

$$\Rightarrow P = \left[\frac{M \cdot q^2 a^2}{6\pi\epsilon_0 c} \right] \rightarrow \frac{M \cdot q^2 \gamma^3 c}{6\pi\epsilon_0} \left(a^2 - \left| \frac{\vec{v} \times \vec{a}}{c} \right|^2 \right)$$

holds for $v \ll c$

cap cond eq.

Liénard's GEN for \vec{v} general

RADIATION REACTION

ACCELERATING CHARGE → RADIATION

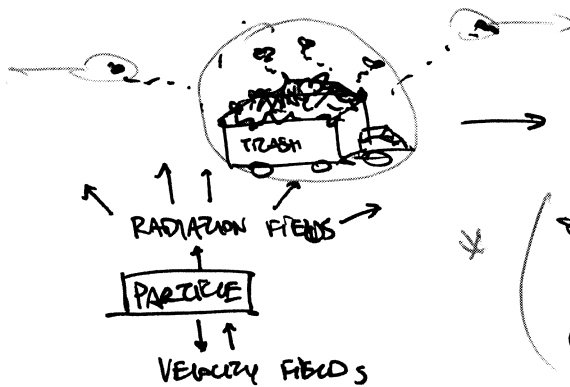
• E CARRIED AWAY @ EXPENSE OF KE (!)

⇒ APPARENTLY RADIATION EXERCISES RECOIL FORCE!

$$P = \frac{M \cdot g^2 q^2}{6\pi\epsilon_0 c^3} \xrightarrow{\text{SKEEVS}} \vec{F}_{\text{rad}} \cdot \vec{v} = - \frac{M \cdot g^2 q^2}{6\pi\epsilon_0 c^3}$$

↑
 $\frac{d}{dt}(\text{force} \cdot \text{dist})$

BUT: ONLY ACCOUNTS FOR RADIATION FIELDS



SO OUR SUGG. IS CORRECT WHEN INTEG OVER FULL CIRCLES (i.e. vel field E GIBES BACK F r.t.u)

TO CALC. RECOIL FORCE, WE REALLY NEED TOTAL POWER LOST AT AN INSTANT (not just tot power that → ∞)

$$E = E^2 = \underbrace{E_v^2}_{\text{"velocity"}} + 2\vec{E}_v \cdot \vec{E}_a + E_a^2$$

$$\int_{t_1}^{t_2} \vec{F}_{\text{rad}} \cdot \vec{v} dt = - \frac{M \cdot g^2}{6\pi\epsilon_0 c^3} \int_{t_1}^{t_2} \dot{q}^2 dt$$

$$\Downarrow \quad \frac{d\vec{v}}{dt} \cdot \frac{d\vec{v}}{dt} dt = \left(v \frac{dv}{dt} \right) \Big|_{t_1}^{t_2} - \int_{t_1}^{t_2} \frac{d^2 v}{dt^2} dt = 0 \text{ BY CIRCLES}$$

$$\int_{t_1}^{t_2} \left(\vec{F}_{\text{rad}} - \frac{M \cdot g^2}{6\pi\epsilon_0 c^3} \ddot{\vec{a}} \right) \cdot \vec{v} dt = 0$$

$$\Rightarrow \boxed{\vec{F}_{\text{rad}} = \frac{M \cdot g^2}{6\pi\epsilon_0 c^3} \dot{\vec{a}}}$$

not a pf!
THE AVG OF ||v|| COMPONENT

STR & WHY WE CARE

• MANY HINTS SO FAR THAT JEP BUILT INTO EIM
 ↳ DEPEND OF LIGHT
 EFFECTIVE LORENTZ CONTRACTION
 EVEN PERIODIC POTENTIAL

eg.

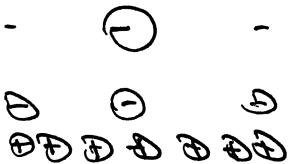
+ - + - + -
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+ → ← + → ← + → ←
 + → ← + → ← + → ←

⊕ wave right
 ⊙ wave left

LAB FRAME: WIRES ATTRACT
 (expl: B-field)

IN rest frame of a - charge in wire 1
 ⇒ Lorentz contraction of + charges in 2



NOT ENERGY FORCE

→ AWESOME .

VELOCITY

VELOCITY IS A TRICKY THING IN SR SINCE

$$V = \frac{d(\text{POSITION})}{d(\text{TIME})}$$

BUT WHICH POSITION & TIME DO WE USE? (which frame)

MOTIVATION: WANT TO USE A VELOCITY THAT IS A 4-VECTOR!

SO CAN USE dx^M IN NUMERATOR
→ DENOMINATOR SHOULD BE AN INVARIANT (ie DOES NOT CHANGE UNDER Λ)

↑ why? b/c dx^M ALREADY HAS THE CORRECT TRANSF.

A GOOD INVARIANT IS THE PROPER TIME $d\tau$
ie. THE TIME YOU MEASURE IN THE OBJECT'S REST FRAME

PROPER VELOCITY $\eta^M = \frac{dx^M}{d\tau}$ $\eta^0 = \frac{cdt}{dz} = \frac{c}{\sqrt{1-u^2/c^2}}$

$$\vec{\eta} = \frac{1}{\sqrt{1-u^2/c^2}} \vec{u} \quad \text{for "ORDINARY VELOCITY" } \vec{u}$$

↑
MAKES SENSE, W/RT SINGLE OBS.

η TRANSF AS A 4 VECTOR: $\eta^M = \Lambda^M_{\nu} \eta^{\nu}$
 \vec{u} TRANSFORMS LIKE A 3-VECT.

Q: WHAT IS $(\eta^M)^2$? c^2 ; WE MOVE IN SPACETIME @ SPD OF LIGHT

MOMENTUM

AGAIN, WANT A 4-VECTOR
NATURALLY $p^M = m\eta^M$

$$p^0 = \frac{mc}{\sqrt{1-u^2/c^2}}$$

IN NATURAL UNITS, $p^M = (E, \vec{p})$ ($c=1$)

$p \cdot p = p^M p_M$ IS A SCALAR (INVARIANT)

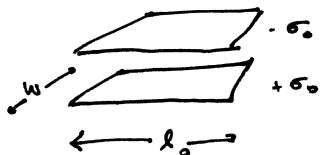
$$\rightarrow \boxed{E^2 - p^2 c^2 = m^2 c^4} \quad \leftarrow \text{how do you get } E = mc^2?$$

↑ CALCULATE @ REST
↑ MUST BE THE SAME
IN ALL FRAMES

HOW FIELDS TRANSFORM

- ASSUME CHARGE IS INVARIANT (KINDA MAKES YOU WONDER WHERE CH. CAME FROM ... GAUGE FREEDOM)
- APPROACH: STUDY TRANSF OF CONST \vec{E} & \vec{B} FIELDS & TRY TO GENERALIZE

CAPACITOR

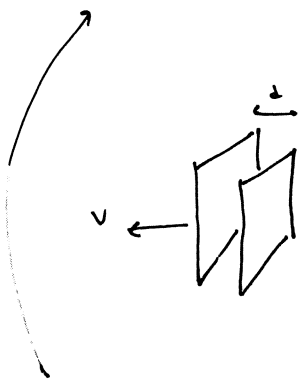


$$\vec{E}_0 = \frac{\sigma_0}{\epsilon_0} \hat{z}$$

GIVE PLATES VELOCITY V GOING LEFT
 $\vec{E}' = \frac{\sigma}{\epsilon_0} \hat{z}$ ONLY DIFF IS σ !
 WHY? GAUSS' LAW

$$\sigma = \gamma \cdot \sigma_0$$

$$\vec{E}'_{\perp} = \gamma \cdot \vec{E}_{\perp}$$



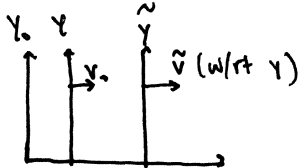
FOR PARALLEL COMPONENT
 E'' INDEP OF d IN CONFIG LEFT.
 $\Rightarrow E'' = E''$

$$\vec{K}'_{\perp} = \gamma \sigma v_0 \hat{x}$$

\Rightarrow

$$\vec{B}'_{\parallel} = -\mu_0 \sigma v_0$$

$$\vec{B}'_{\parallel} = -\mu_0 \tilde{\sigma} \tilde{v}$$



$$\tilde{v} = \frac{v + v_0}{1 + vv_0/c^2}$$

$$\tilde{\sigma} = \tilde{\gamma} \sigma$$



$$\vec{B} = \gamma \mu_0 \vec{I} \quad \text{? so Arth.}$$

Math \Rightarrow

$$\vec{E}'_{\perp} = \gamma (\vec{E}_{\perp} - v \vec{B}_{\parallel})$$

$$\vec{B}'_{\perp} = \gamma (\vec{B}_{\perp} + \frac{v}{c^2} \vec{E}_{\parallel})$$

$$\vec{E}'_{\parallel} = E_{\parallel}$$

$$\vec{B}'_{\parallel} = B_{\parallel}$$

WE NOTICE THIS HAS THE SAME TRANSFORMATION PROPERTIES AS A RANK 2 ANTI-SYMMETRIC TENSOR

$$\hookrightarrow t^{\mu\nu} = -t^{\nu\mu}$$

$$t^{\mu\nu} \rightarrow \Lambda^{\mu}_{\alpha} \Lambda^{\nu}_{\beta} t^{\alpha\beta}$$

RELATIVISTIC E & M

4-VECTOR CURRENT DENSITY: $\boxed{j^M = \rho_0 \eta^M} = (c\rho, \vec{j})$

CONTINUITY EQ: $\nabla \cdot \vec{j} = -\dot{\rho} \Leftrightarrow \boxed{\partial_\mu j^M = 0}$

Maxwell: $\boxed{\partial_\nu F^{\mu\nu} = \mu_0 j^\mu, \quad \partial_\nu G^{\mu\nu} = 0}$

$$F^{\mu\nu} = \begin{pmatrix} 0 & E/c \\ -E/c & 0 & B_3 - B_1 \\ B_1 & 0 & B_2 \\ B_1 - B_2 & 0 & 0 \end{pmatrix}$$

$$G^{\mu\nu} = F^{\mu\nu} \omega \quad \begin{matrix} E/c \rightarrow \vec{B} \\ \vec{B} \rightarrow -E/c \end{matrix}$$

FORCE LAW MINKOWSKI FORCE $\boxed{K^M = \int \eta_{\nu\mu} F^{\mu\nu}}$

$$\frac{dP^M}{d\tau} = \frac{dt}{d\tau} \frac{dP^M}{dt} \Rightarrow \vec{K} = \frac{1}{\sqrt{1-u^2/c^2}} \vec{F}$$

A BETTER FORMULATION: POTENTIALS

$$A^M = \left(\frac{V}{c}, \vec{A} \right)$$

$$F^{\mu\nu} = \partial_\mu A^\nu - \partial_\nu A^\mu$$

SIMPLEST ANTSYMMETRIC
2 TENSOR WE COULD MAKE
OUT OF DERIVATIVES OF A

MAXWELL: $\boxed{\square^2 A^M = -\mu_0 j^M}$ (LORENZ GAUGE)

IF WE WERE CLEVER, WE COULD HAVE GUESSED THIS!

$$\begin{aligned} \vec{B} &= \nabla \times \vec{A} & \Rightarrow & B_2 = \partial_1 A^3 - \partial_3 A^1 \\ \vec{E} &= -\nabla V - \frac{\partial \vec{A}}{\partial t} & \Rightarrow & \frac{1}{c} E_2 = \partial_3 A^0 - \partial_0 A^3 \end{aligned}$$