

Cornell Laboratory for Accelerator-based Sciences and Education (CLASSE)

# First Cavity Results from the Cornell SRF Group's Nb<sub>3</sub>Sn Program

NA PAC'13

\*Best\* Wuppertal Cavity, 2.0 K
 \*Best\* Wuppertal Cavity, 4.2 K
 Cornell ERL1-4, 2.0 K
 Cornell ERL1-4, 4.2 K

 $\land \land \land$ 

10

Sam Posen and Matthias Liepe Cornell University Thursday, Oct 3, 2013 NA-PAC '13, Pasadena, California

Eacc [MV/m



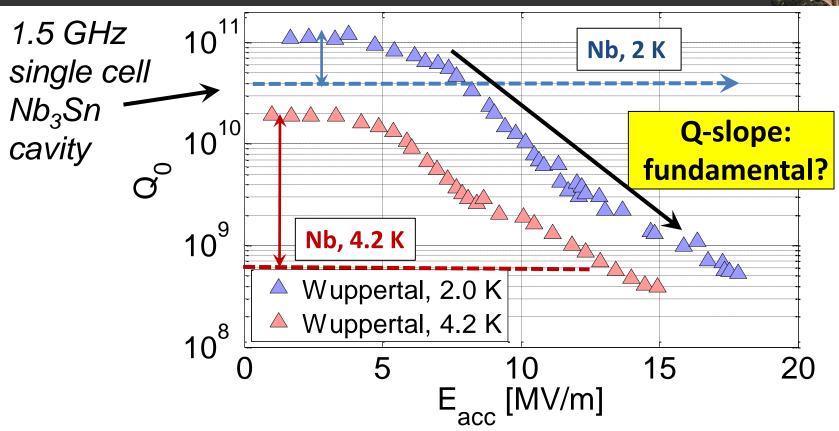


- Nb<sub>3</sub>Sn has T<sub>c</sub> of ~18 K, vs ~9 K for Nb: much lower BCS R<sub>s</sub>(T)
- Significant reduction in losses at same temperature
- Possibility to operate at higher temperatures: *LHe at atmospheric pressure? Cold gas?*
- Smaller cryo plant and less grid power
  - Application to CW SRF linacs for light sources, small scale accelerators (closed He gas cryogenic system for universities/hospitals), industrial applications (wastewater and flue gas treatment, isotope production)
- Higher predicted superheating field ~400 mT, nearly twice Nb
  Application to high energy SRF linacs: reduce # of cavities



#### Previous SRF Research with Nb<sub>3</sub>Sn

NARACI 3. LBL.GOV PAC '13



- Excellent R<sub>s</sub> at low fields, but large increase in R<sub>s</sub> with field ("Q-slope") above ~5 MV/m
- Various suggested causes: intergrain losses, bad stoichiometry, and vortex penetration at lower critical field B<sub>c1</sub>

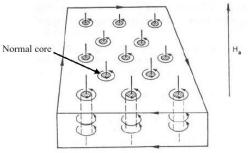




 $B_{c1}$  is the onset of metastability. Above  $B_{c1}$ , an energy barrier prevents vortex penetration, but surface defects of size  $\sim \xi$  lower barrier.

#### Is $\xi$ of Nb<sub>3</sub>Sn so small that B<sub>c1</sub> is the limit?

ξ of Nb ~ 20-30 nm ξ of Nb<sub>3</sub>Sn, NbN, MgB<sub>2</sub> ~ 3-4 nm



# If vortices penetrate at B<sub>c1</sub>, **all** alternative SRF materials would be severely limited.

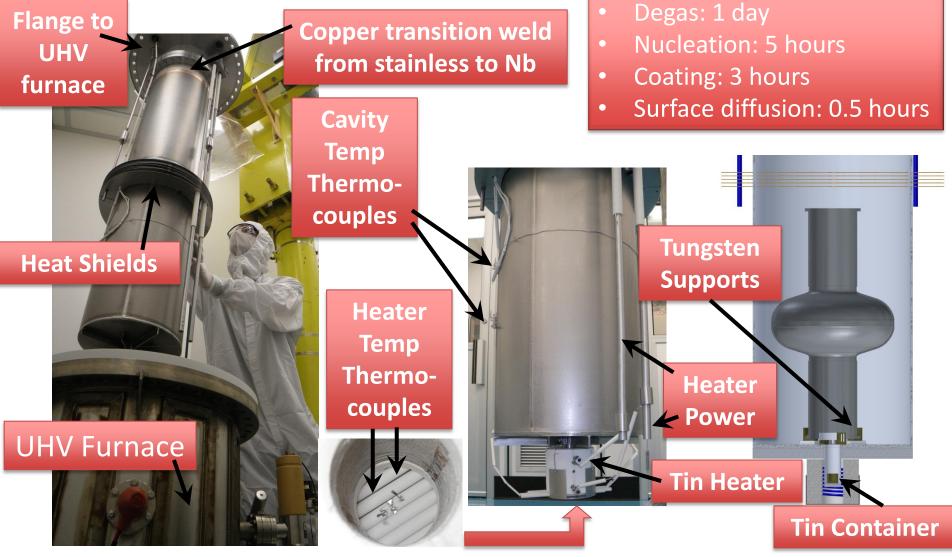






#### **Cornell Cavity Coating Chamber**





See THPO066, SRF11 for details of coating process and commissioning process using samples.

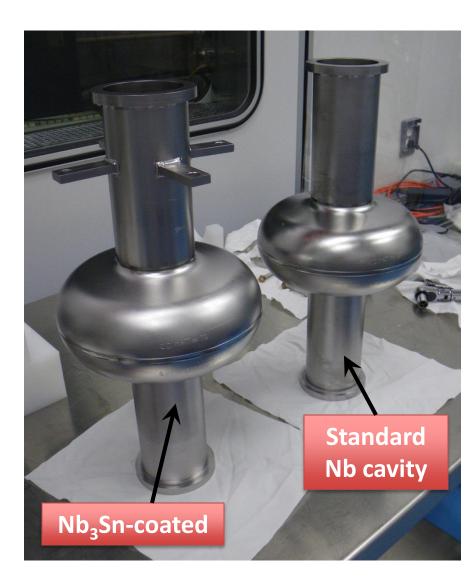






### Cornell Nb<sub>3</sub>Sn Coated Cavity



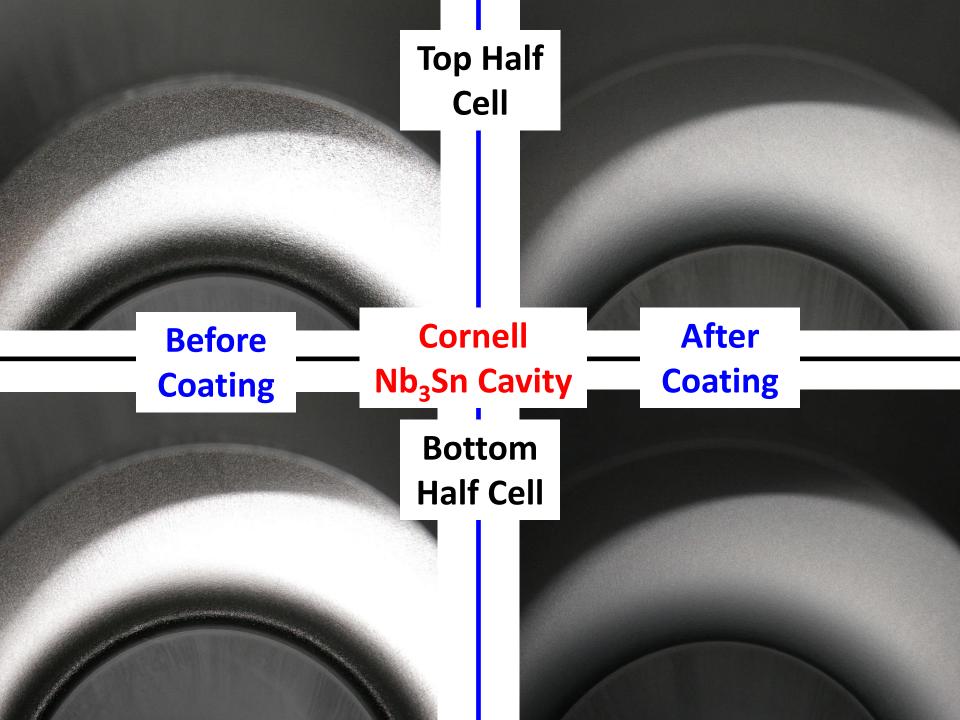




S. Posen - First Cavity Results from the Cornell Nb<sub>3</sub>Sn Program



NAC PAC'13



# Breakthrough Nb<sub>3</sub>Sn Cavity

- New Nb<sub>3</sub>Sn cavity: ERL shape (similar to TESLA), single cell, 1.3 GHz
- Tested after very slow cool (>~6 min/K)
- Excellent performance, especially at 4.2 K
- The first accelerator cavity made with an alternative superconductor that far outperforms Nb at usable gradients





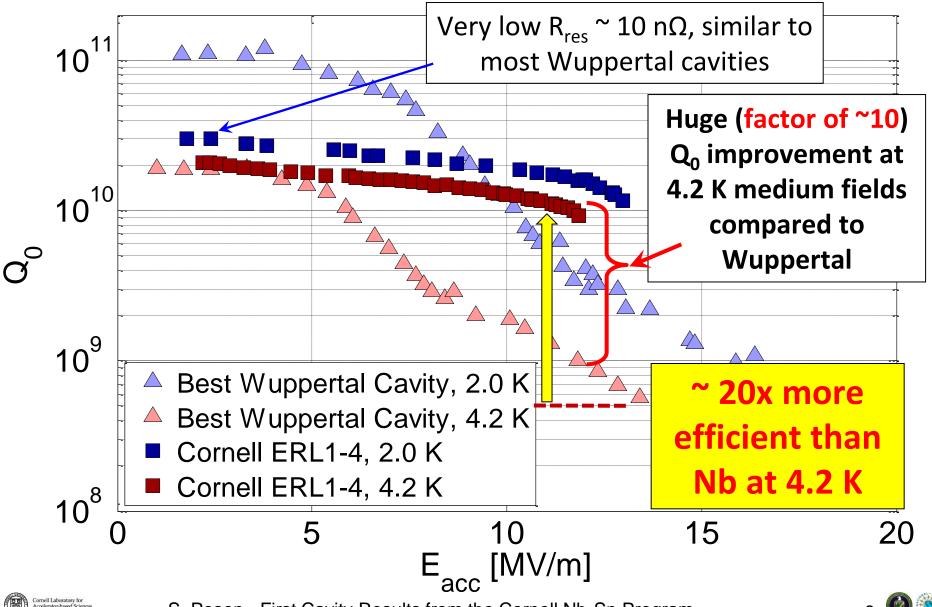


PAC '13

# Breakthrough Nb<sub>3</sub>Sn Cavity

13

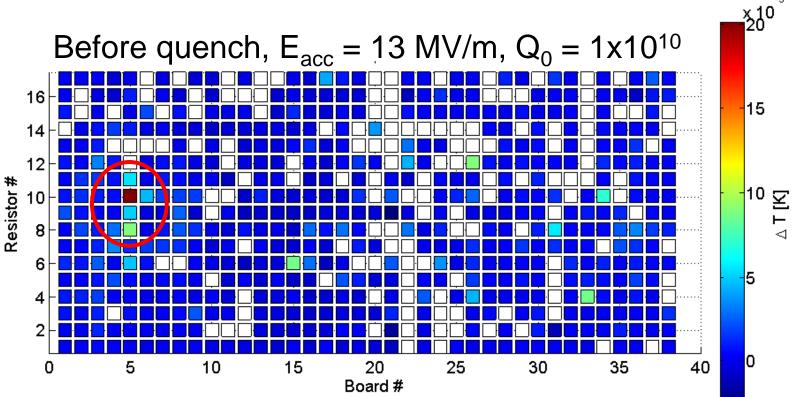
PAC







- Localized pre-heating just below first quench
- Defect not a fundamental limit
- Can reach higher fields by fixing defect

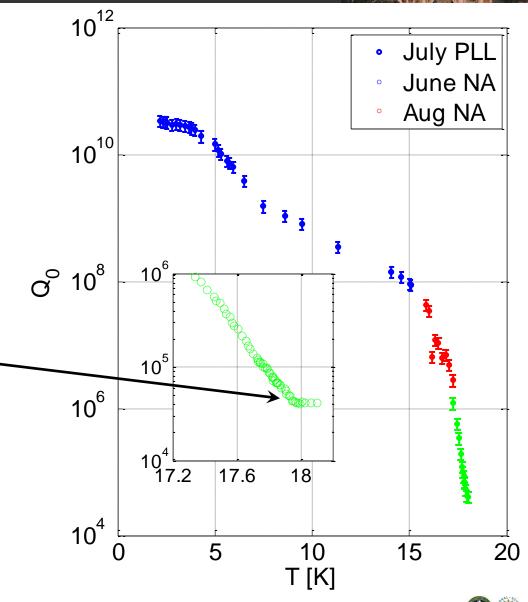








- No sign of Q<sub>0</sub> change near T<sub>c</sub> of niobium: excellent Nb<sub>3</sub>Sn coverage!
- High T<sub>c</sub> of 18.0 K close to maximum literature value
- Extract material parameters from this data



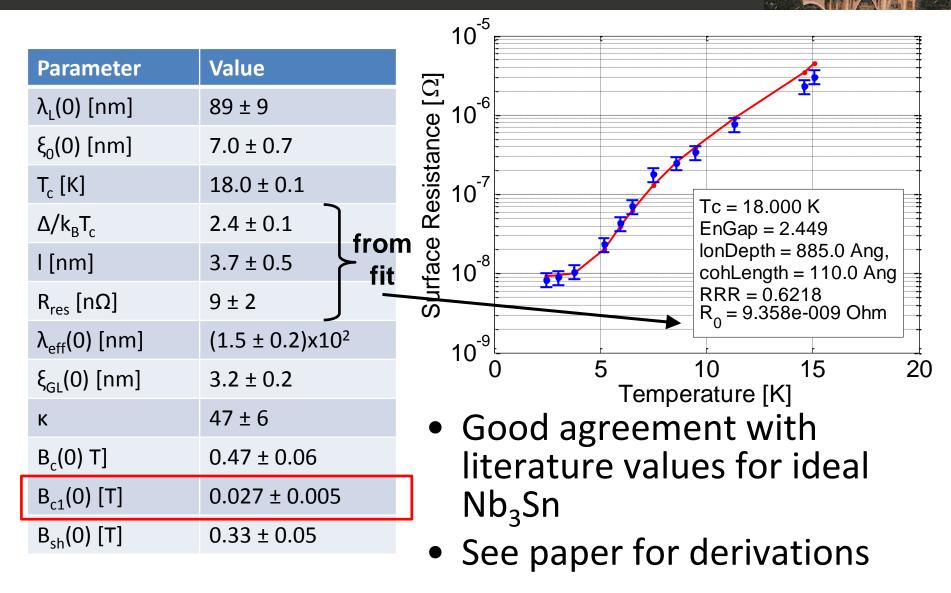
13

PAC





## Fits to Material Parameters







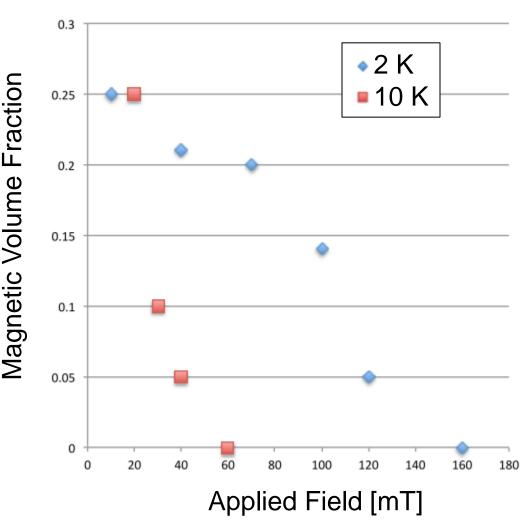
**'13** 

PAC



# Sample B<sub>c1</sub> Measurement

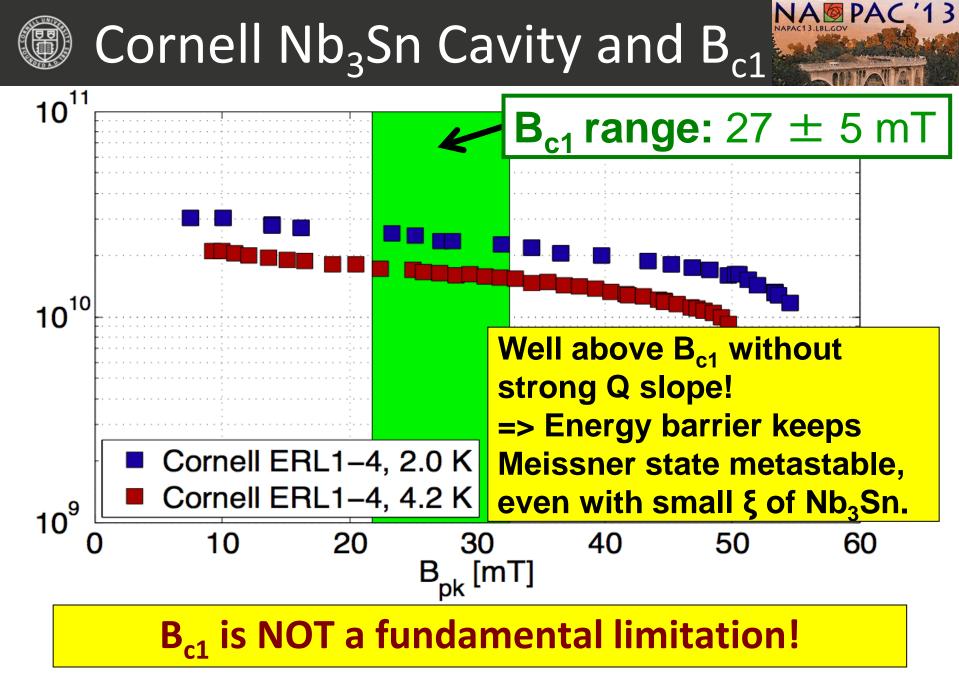
- B<sub>c1</sub> of Nb<sub>3</sub>Sn sample measured directly via μ-SR by Anna Grassellino et al.
- B<sub>c1</sub> ~ 20-30 mT agrees with cavity measurement



A. Grassellino et al. , TUP029 (Presented at the SRF Conference, Paris, France, 2013).













# Conclusions



 Current status: Nb<sub>3</sub>Sn is now a promising alternative SRF material for certain future accelerators: Cornell cavity demonstrated at 4.2 K, usable gradients ~12 MV/m, Q<sub>0</sub> of 10<sup>10</sup>, 20 times higher than Nb

Q-slope seen in previous cavities not a fundamental problem

- Near future: fix high performing but defect limited cavity, or coat new one—expect even higher gradients
- Longer Term R&D Plan: Develop surface preparation methods for Nb<sub>3</sub>Sn to push performance (as has been done in Nb over many years)
- Eventual Hope: Prevent non-fundamental limitations to reach fields close to ultimitate limit, B<sub>sh</sub> ~ 400 mT



